

Models for Global Plasma Dynamics

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Outline

1 Models for Long-Wavelength Plasma Dynamics

- Introduction
- Fluid models: MHD and Hall-MHD
- Gyrokinetic theory

2 Hamiltonian Gyrofluid Reconnection

- Motivation
- HEMGF: A Hamiltonian Gyrofluid Model
- Reconnection

Outline

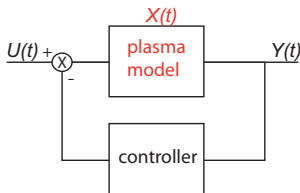
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Relationship to control design process



$$\dot{X}(t) = f(X(t), U(t), t);$$
$$Y(t) = g(X(t), U(t), t).$$

- We focus on first two elements of control design process
 - 1 Make system model
 - 2 Verify model predicts behavior of system
 - 3 Design controller
 - 4 Test models in closed-loop simulation
 - 5 Implement and test implementation
 - 6 Deploy in operation
- We are concerned in this and subsequent talks with formulating and solving the equation for $X(t)$.

State space and Equilibrium

- Plasmas have infinite degrees of freedom, so X is a denumerably infinite-dimensional vector. We can think of it in terms of its Fourier coefficients. To represent X on a computer it must be *truncated*. Understanding the effect of truncation is part of the modeling task.
- Think of plasma as a system close to a quiescent, equilibrium state X_0 :

$$\dot{X}_0 = F(X_0, U, t) = 0$$

Note that $X_0 = X_0(U)$.

Linear solution

- The singular value decomposition leads to the solution:

$$X(t) = X_0 + \sum e^{i\omega_j t} |v_j\rangle \langle u_j|.$$

- In practice, nonlinear effects are often important.
- To develop a useful model, we need to eliminate irrelevant time scales: for designing an steam engine, plate tectonics and nuclear vibrations are irrelevant.

Time scales I: The centrifugal governor



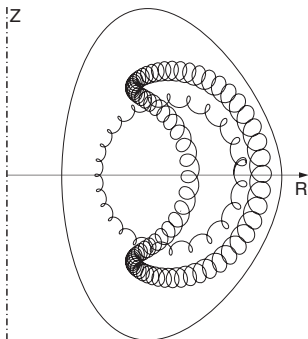
- T. Mead (1787), J.C. Maxwell (1867).

$$\begin{aligned}\ddot{\theta} &= \omega^2 \sin \theta \cos \theta - (g/\ell) \sin \theta; \\ l\dot{\omega} &= T(\theta).\end{aligned}$$

- For a stable governor, we can eliminate the oscillatory degree of freedom: $\cos \theta = g/\ell\omega^2$.
- Alternatively, use implicit methods:

$$\begin{aligned}X^{j+1} &= X^j + hF(X^{j+1}) \\ X^{j+1} &= X^j + h(1 - h\mathbf{M})^{-1}F(X^j).\end{aligned}$$

Time scales (II): single particle motion



- Single particle motion in a tokamak exhibits five fundamental time-scales:
 - 1 The electron and ion cyclotron period

$$\tau_{cs} = m_s / e_s B$$

- 2 The electron and ion transit times

$$\tau_{ts} = R / v_{ts} = (R / \rho) \tau_{cs}$$

- 3 The drift time

$$\tau_D = (\rho_s / R) \omega_{ts} \sim T / e B R^2$$

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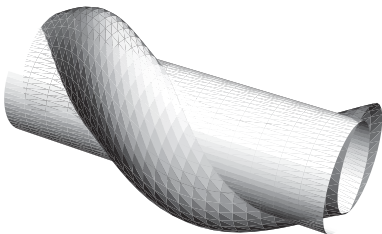
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MHD assumes the plasma motion is kink-driven



- The total momentum conservation law is:

$$m_i n \frac{d\mathbf{V}}{dt} = -\nabla p - \nabla \cdot \Pi + \mathbf{J} \times \mathbf{B},$$

where $\mathbf{V} = \mathbf{V}_i$.

- Balancing inertia against the kink force determines the time-scale

$$\tau_A = R/V_A = \beta_i^{1/2} \tau_{ti}.$$

Electron momentum conservation determines the electric field

- Using $\mathbf{V}_e = \mathbf{V} - \mathbf{J}/ne$ in the magnetic force, the electron momentum conservation takes the form

$$m_e n \frac{d\mathbf{V}_e}{dt} = ne(\mathbf{E} + \mathbf{V} \times \mathbf{B} - \eta \mathbf{J}) - \nabla p_e - \nabla \cdot \Pi_e - \mathbf{J} \times \mathbf{B}$$

- The only term that can balance the magnetic force is thus the electric force:

$$\mathbf{E} + \mathbf{V} \times \mathbf{B} = 0.$$

Digression: Single vs. Two-fluid models

- MHD is sometimes called a “single fluid” model. But plasma conducts electricity: $\mathbf{J} = ne(\mathbf{V}_i - \mathbf{V}_e) \neq 0$.
- It is easy to see that

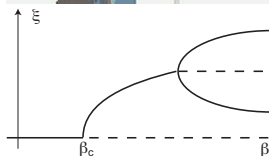
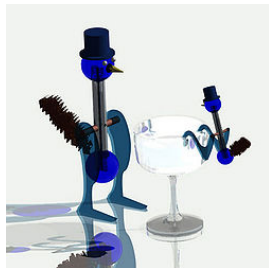
$$\frac{V_i - V_e}{V_i} \sim \frac{J}{neV_A} \sim \frac{d_i}{a} = \rho_* \beta_i^{-1/2} \ll 1,$$

where $\rho_* = \rho_i/a$.

- So MHD is a “quasi-single-fluid” in the same sense that it is a “quasi-neutral” theory:

$$\frac{n_i - n_e}{n_e} = \frac{\epsilon_0 \nabla \cdot \mathbf{E}}{n_e e} = \left(\frac{V_A}{c} \right)^2 \rho_* \beta_i^{-1/2} \ll 1.$$

Fusion MHD events are sub-Alfvénic



- Typical times scales for sawtooth crash, ELM, disruption $\sim 100\mu s \ll \tau_A \sim 1\mu s$.
- In the drinking bird toy, evaporation draws fluid up the tube, creating an inverted pendulum (c.f. D. Humphreys lecture)
- The evaporation rate $\ll \gamma$, so the dip is preceded by a precursor oscillation.
- Fusion plasmas do not fit the “drinking bird” paradigm: instabilities
 - 1 often lack a discernible precursor.
 - 2 Never satisfy ideal MHD.

Sub-Alfvenic MHD motion exhibits spatial resonances

- For $V \ll V_A$, the MHD equilibrium equation applies

$$\mathbf{J} \times \mathbf{B} = \nabla p.$$

- This determines $\mathbf{J}_\perp = (\mathbf{B} \times \nabla p)/B^2$.
- Ampere's law $\nabla \cdot \mathbf{J} = 0$ determines $J_\parallel$:

$$\mathbf{B} \cdot \nabla (J_{\parallel}/B) = -\nabla \cdot [(\mathbf{B} \times \nabla p)/B^2].$$

- The operator $\mathbf{B} \cdot \nabla = (m - nq)B/Rq$ is **singular** on magnetic surfaces with closed field lines.

Sub-Alfvénic motions are described by the drift ordering

- Recall the electron momentum conservation

$$m_e n \frac{d\mathbf{V}_e}{dt} = ne(\mathbf{E} + \mathbf{V} \times \mathbf{B} - \eta \mathbf{J}) - \nabla p_e - \nabla \cdot \Pi_e - \mathbf{J} \times \mathbf{B}$$

- Assume $\mathbf{V} \times \mathbf{B} \sim \nabla p_e / ne$ This is the *drift ordering*.
- Eliminate $\mathbf{E}$ between the electron momentum conservation and Faraday's law,

$$\frac{\partial \mathbf{B}}{\partial t} = -\nabla \times \mathbf{E}.$$

Hall Magnetohydrodynamics

- There follows

$$\frac{\partial \hat{\mathbf{B}}}{\partial t} = -\nabla \times \hat{\mathbf{E}},$$

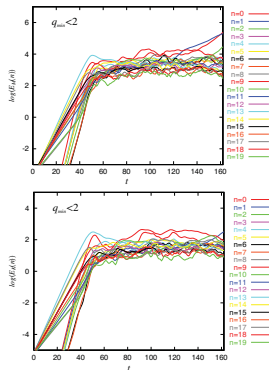
where

$$\hat{\mathbf{E}} = \mathbf{V} \times \mathbf{B} - \mathbf{J} \times \hat{\mathbf{B}}/ne - (\nabla p_e + \nabla \cdot \Pi_e)/ne;$$

$$\hat{\mathbf{B}} = (1 - d_e^2 \nabla^2) \mathbf{B};$$

$$\mathbf{J} = \nabla \times \mathbf{B} / \mu_0.$$

The perils of two-fluid models



- 1 Hall MHD describes whistler waves and the ion cyclotron resonance. This is the **curse of the drift ordering**: to describe slow evolution we have to include fast waves!

$$\omega = kV_A \rightarrow \omega = k^2 \rho_s V_A$$

- 2 By including drift motion, two-fluid models enable all the drift instabilities!
- 3 The model is still missing important kinetic physics (FLR, parallel dynamics, etc. . .)
- 4 It is hard to parallelize.

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Brief outline of gyrokinetic theory

- The GKM is based on an expansion in $\rho_* = \rho_i/a$ such that the distribution function is allowed to exhibit rapid variations across the field, but with *small amplitude* $F = F_0 + F_1$ where $F_1 = O(\rho_*)$ but $\nabla F_1 \sim \nabla F_0$.
- The equation is

$$\frac{\partial F}{\partial t} + (v_{\parallel} \mathbf{b} + \mathbf{v}_E + \mathbf{v}_D) \cdot \nabla F + \left[\frac{e}{m} E_{\parallel} - \mu \mathbf{b} \cdot \nabla B + v_{\parallel} (\mathbf{b} \cdot \nabla \mathbf{b}) \cdot \mathbf{v}_E \right] \frac{\partial F}{\partial v} = 0.$$

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The Gyrofluid Model (GFM): I. Moments

The gyrokinetic equations are computationally intensive. Taking their moments yields a more manageable fluid model:

$$\frac{dn}{dt} + \bar{\mathbf{b}} \cdot \nabla u_{\parallel} + \frac{1}{2}(\hat{\nabla}_{\perp}^2 \bar{\mathbf{v}}) \cdot \nabla T_{\perp} + \dots = 0; \quad (1)$$

$$\frac{dP}{dt} + \bar{\mathbf{b}} \cdot \nabla p - \frac{1}{2}(\hat{\nabla}_{\perp}^2 \bar{\mathbf{b}}) \cdot \nabla T_{\perp} + \dots = 0, \quad (2)$$

where

$$\bar{\mathbf{b}} = (\mathbf{B}_0 + \mathbf{B}_0 \times \nabla \Psi) / B_0, \quad \frac{d}{dt} = \frac{\partial}{\partial t} + \bar{\mathbf{v}} \cdot \nabla, \quad \bar{\mathbf{v}} = \mathbf{b}_0 \times \nabla \Phi,$$

and $\Psi = \Gamma_0^{1/2} \psi$, $\Phi = \Gamma_0^{1/2} \phi$ are the gyro-averaged magnetic and electrostatic potentials.

GFM II. Gyro-averaging

Dorland and Hammett defined the operator $\Gamma_0^{1/2}$ as follows

$$\Gamma_0^{1/2}\xi = \exp(\frac{1}{2}\tau\nabla_\perp^2)I_0^{1/2}(-\tau\nabla_\perp^2)\xi, \quad (3)$$

where $\tau = T_s/T_{\text{ref}}$ and I_0 is the modified Bessel function of the first kind. The definition in Eq. (3) should be interpreted in terms of its series expansion

$$\Gamma_0^{1/2}\xi = 1 + \sum_{n=1}^{\infty} a_n(\tau\nabla_\perp^2)^n = 1 + (\tau/2)\nabla^2 + \dots,$$

where the a_n are real numbers.

GFM III. Ampère and quasi-neutrality

- The system is completed by
 - 1 the quasi-neutrality equation

$$n_e = \Gamma_0^{1/2} n_i + (\Gamma_0 - 1) \phi / \tau,$$

- ## 2 Ampère's law

$$J = \nabla^2 \psi = -\frac{\tau \beta_e}{2} (\Gamma_0^{1/2} u_i - u_e).$$

- The GFM experienced a period of success before being discredited due its overestimation of zonal flow damping.
- Its electromagnetic version is coming back into use in space physics and for studies of magnetic reconnection.

Hamiltonian versions of the GFM model can be constructed

- Why Hamiltonian? Recall that in MHD the equilibrium and charge conservation equations require that

$$\mathbf{B} \cdot \nabla (J_{\parallel}/B) = -\nabla \cdot (\mathbf{B} \times \nabla P/B^2).$$

- Integrability of the charge conservation condition requires that

$$\oint \frac{d\ell}{B} \nabla \cdot \mathbf{J}_\perp = 0.$$

- Similar constraints must be satisfied by the vorticity, density, etc. . . **When are these conditions satisfied?**

Hamiltonians at work

- The ideal part of the fluid equations must be expressible in terms of Poisson brackets:

$$\partial_t \xi_j = \{\xi_j, H\} + D \nabla^2 \xi_j.$$

- Poisson brackets generally possess families of geometrical invariants C_k called Casimirs.

$$\{\xi_j, C_k\} = 0.$$

- Equilibria are the extrema of the functional $F = H + \sum_k C_k$,

$$\delta F = 0.$$

The Hamiltonian formulation thus guarantees equilibrium integrability.

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Hamiltonian formulation for Alfvén dynamics

- Consider the following model that describes kinetic and inertial-Alfvén waves:

$$\frac{\partial n_i}{\partial t} + [\Phi, n_i] = 0; \quad (4)$$

$$\frac{\partial n_e}{\partial t} + [\phi, n_e] - c_A^2 \nabla_{\parallel} J = 0; \quad (5)$$

$$\frac{\partial}{\partial t}(\psi - d_e^2 J) + [\phi, \psi - d_e^2 J] + \nabla_{\parallel} n_e = 0, \quad (6)$$

where

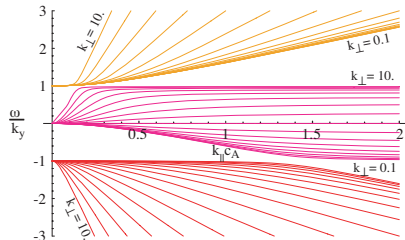
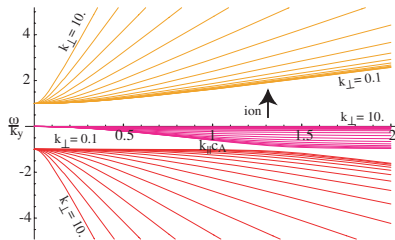
$$[f, g] = \mathbf{B}_0 \cdot (\nabla f \times \nabla g) / B_0.$$

- The conserved energy is

$$H = \frac{1}{2} \langle c_A^2 (|\nabla \psi|^2 + d_e^2 J^2) + n_e^2 + \Phi n_i - \phi n_e \rangle.$$

Dispersion relation

- The gyrofluid equation reproduces *exactly* the kinetic dispersion relation in both the kinetic-Alfvén and inertial regimes.
- Unlike FLR (Braginskii) models, it reproduces the band gap between the ion and electric drift frequencies.



Normal fields

- The Casimirs suggest the use of the normal fields n_i and $G_{\pm} = \psi - d_e^2 \nabla^2 \psi \pm \rho_e n_e$.
- The equations of motion for the normal fields are

$$\frac{\partial n_i}{\partial t} + [\Phi, n_i] = 0; \quad (7)$$

$$\frac{\partial G_{\pm}}{\partial t} + [\phi_{\pm}, G_{\pm}] = 0; \quad (8)$$

$$(9)$$

where

$$\phi_{\pm} = \phi \pm \psi / \rho_e$$

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Without magnetic reconnection, MHD instabilities would be harmless

- The “frozen-in” property of ideal MHD means that plasma cannot cross flux surfaces, and flux surfaces cannot break, so an ideally unstable plasma would either settle into a bifurcated state or bounce on the walls like a balloon.
- Magnetic reconnection prevents the saturation of MHD instabilities. The conditions for the onset of magnetic reconnection depend on plasma flows (diamagnetic and electric), FLR effects, etc. . .
- The onset conditions are incompletely understood.

Role of ion temperature in reconnection

- For constant-density, the model reduces to two fields:

$$\frac{\partial \mathbf{G}_{\pm}}{\partial t} + [\phi_{\pm}, \mathbf{G}_{\pm}] = 0,$$

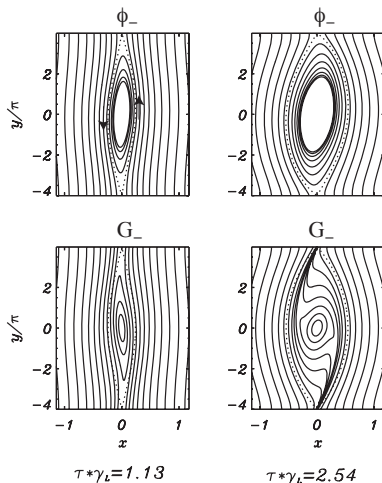
where $G_{\pm} = \psi - d_e^2 \nabla^2 \psi \pm \rho_e (1 - \Gamma_0) \phi$

- Solving for ψ and ϕ yields

$$\begin{aligned}\psi &= \frac{1}{2}(1 - d_e^2 \nabla^2)^{-1}(G_+ + G_-); \\ \phi &= \frac{1}{2\rho_e}(1 - \Gamma_0)^{-1}(G_+ - G_-).\end{aligned}$$

- We compare the cases of cold and hot ions

Phase-mixing in cold-ion reconnection



- The Lagrangian quantities $G_{\pm}$ are convected by the $\phi_{\pm}$ stream-functions.
- For cold ions,

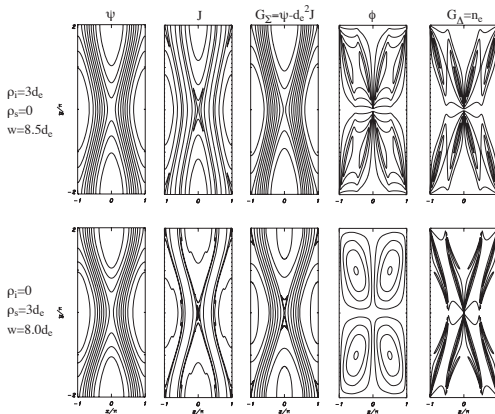
$$(1 - \Gamma_0)^{-1} = \nabla^{-2},$$

so ϕ and ψ are smoothed.

- This is analogous to Landau damping.

(Grasso, Califano, Pegoraro and Porcelli, PRL 2001.)

Gyrofluid reconnection



- For hot ions, by contrast,

$$(1 - \Gamma_0)^{-1} = -\rho_i^2 + \nabla^{-2}$$

so ϕ is **not** smoothed
(Grasso, Califano, Pegoraro
and Porcelli, Plasma Phys.
Reports 2000.)

Summary (I)

- MHD is inadequate to model most MHD events in tokamaks.
- Two-fluid models provide a better description, but they still neglect important effects such as parallel heat flow, Landau damping, and finite Larmor radius.
- Two-fluid models unleash onto MHD the pandora's box of drift-acoustic turbulence: ITG, ETG, etc. . . Subgrid models may need to be developed.

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Summary (II)

- Gyrokinetic theory is free from the deficiencies of fluid models. Two of its own deficiencies are subjects of ongoing research:
 - 1 Quasi-static coupling to the compressional Alfvén wave;
 - 2 Collision operators
- The gyrofluid closure method holds considerable promise for understanding the role of FLR in MHD.
- The model that will enable us to understand ITER MHD has yet to be invented, perhaps by one of you.

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